# Laboratory formation of fullerenes from PAHs: Top-down interstellar chemistry.

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- $\bullet~C_{60}$  (and  $C_{70})$  are the largest molecules detected in the ISM.
- Highly stable molecule.
- Its formation is related to other carbon bearing compounds.
- Ubiquitous in the ISM.
- Possible culprit of some DIBs.



## Observations

- First confirmed detection (along with C<sub>70</sub>) in circumstellar environment of PN Tc 1 (Cami et al. 2010).
- Later detections in wide variety of environments and physical conditions: PNe, post-AGB stars, YSOs, Herbig Ae/Be and PDRs associated with RNe and HII regions.



Cami et al. (2010).

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- Formation in the envelope of AGBs has limited efficiency (Bernard-Salas et al. 2012).
- Berné & Tielens (2012) proposed that PAH dehydrogenation, followed by C<sub>2</sub>-losses, can form C<sub>60</sub>.
- Micelotta et al. (2012) propose a similar mechanism, but starting from HAC.



Sellgren et al. (2010).

# Top-down Chemistry



Berné & Tielens (2012).

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Fullerenes and PAHs

PD in Astrochem.



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# "Magic Numbers"





- Formation of fullerenes shows peaks with enhanced intensity.
- "Magic numbers":  $C_{44}$ ,  $C_{50}$  and  $C_{56}$ .



- C<sub>60</sub> has an absorption minimum at  $\sim$ 500–600 nm.
- C<sub>70</sub> and large PAHs absorb efficiently.
- At shorter wavelengths the absorption becomes comparable for the three.



Tatsuhisa et al. (1991), Malloci et al. (2007)

- Fullerenes follow "cage route" only.
- Irradiation with 266, 355 and 532 nm.
- C<sub>60</sub><sup>+</sup> does not dissociate at 532 nm.

#### Goals:

- Dehydrogenation and C<sub>2</sub> loss.
- Compare dissociation patterns.
- C<sub>2</sub> loss necessary for isomerization?



# Dehydrogenation



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## $266 \ {\rm and} \ 355 \ {\rm nm}$ Irradiation



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## 532 nm Irradiation



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# $C_{78}H_{26}$ Fragmentation



- $C_{60}$  can be formed from large PAHs efficiently.
- Large PAHs pass through a first step of fast dehydrogenation.
- C<sub>2</sub> losses are a necessary step for isomerization of closed-cage.
- Smaller (and larger) cages can also be formed.

#### Future Work:

- Confirmation with IR spectroscopy.
- Derive energies involved using synchrotron radiation.